# VARIATIONAL CALCULATIONS ON THE ${ }^{2} \mathbf{P}_{1 / 2}$ GROUND STATE OF BORON ATOM USING HYDROGENLIKE ORBITALS* 

MARIA BELÉN RUIZ AND MIGUEL ROJAS<br>Chair for Theoretical Chemistry of the Friedrich-Alexander University Erlangen-Nürnberg,<br>Egerlandstraße 3, D-91058 Erlangen, Germany<br>e-mail: Maria. Belen. Ruiz@chemie. uni-erlangen.de

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#### Abstract

Variational calculations of the ${ }^{2} \mathrm{P}_{1 / 2}$ ground state of boron atom are performed using a singleterm reference wave function and a 150 -term wave function expansion without interelectronic distances. The wave function is constructed with hydrogenlike orbitals. These orbitals are superior to Slater orbitals, because the orbital 2 s contains nodes. The calculated energy $-24.550233 \mathrm{a} . \mathrm{u}$. is compared with -24.541246 a.u. using Slater orbitals and the same basis function expansion, and with -24.5689998 a.u. obtained from full-CI calculations using a $4-31 \mathrm{G}$ basis set. The single-term wave function constructed with hydrogenlike orbitals leads to an energy value of 24.501187 a.u., which is lower than the HartreeFock energy using a single-zeta basis set of Slater orbitals and it is also lower than with a single-term wave function with Slater orbitals, both lead to an energy of -24.498369 a.u. The behavior of the node of the 2 s orbital and its radial distribution function of the wave function series are discussed.


## 1. INTRODUCTION

Hydrogenlike orbitals contain a polynomial in $r$ while the Slater orbitals are monomial. In addition, the hydrogenlike orbitals have nodes. Hydrogenlike orbitals have been used in calculations of the first row atoms by Zener [1] and Eckart [2]. The energy results were equivalent to the ones obtained with Slater orbitals [3]. Indeed, a hydrogenlike orbital is a linear combination of Slater ones [4]. Other similar orbitals were used by Morse [5] and Tubis [6]. In this article we use a $2 s$ orbital of Morse-type with two adjustable parameters.

The purpose of this work is to use hydrogenlike atomic orbitals to improve the quality of the orbitals of a single-term and many-term wave functions, in order to have a good starting reference wave function for a full-Hylleraas calculation. The Hylleraas method [7] is a variational method which introduces the correlation effects by including explicitly the interelectronic distances in the wave function.

The Hylleraas-type wave functions are linear expansions of basis functions containing interelectronic coordinates whose coefficients are determined variationally. The calculations of the atoms helium [8] (other sets of basis functions have been also used), lithium [9-11, 14], and beryllium [12, 13] yielded energies close to the experiments. Parallel to the Hylleraas

[^0]method, Hylleraas-CI and exponentially correlated wave functions yielded highly accurate energy results for helium [15], lithium [16], beryllium [17-19] and boron [20] atoms.

Recently, calculations on the ground state of boron atom have been made using the singleterm and 150 -term wave functions constructed with Slater orbitals [21]. The obtained energies lie between the Hartree-Fock and the CI energies, including about 28 per cent of the correlation energy. In this paper we use the same wave function but constructed with hydrogenlike orbitals and we compare with the results using Slater ones.

## 2. THEORY

The wave function of the ${ }^{2} P_{1 / 2}$ state of boron atom is written:

$$
\begin{equation*}
\Psi=\hat{\mathrm{A}} \phi \chi \tag{1}
\end{equation*}
$$

where $\phi$ is the spatial function, $\hat{A}$ is the 5 -particle antisymmetrization operator, $\chi$ is the spin function $\alpha \beta \alpha \beta \alpha$. Since it is possible to make the restriction of using the same spatial function fortwo electrons in the same shell, the single-determined wave function is eigenfunction of $\hat{S}^{2}$ with eigenvalue $M_{s}=M_{z}, M_{z}$ being the eigenvalue of $\hat{S}_{z}$. Moreover, the spatial functions may be chosen such that the total wave function is eigenfunction of $\hat{L}^{2}$.

We construct a trial wave function as the expansion [10]

$$
\begin{equation*}
\Psi=\sum_{\mu=1}^{N} C_{\mu} \hat{\mathrm{A}} \phi_{\mu} \chi=\sum_{\mu=1}^{N} C_{\mu} \psi_{\mu} \tag{2}
\end{equation*}
$$

where the constants $C \mu$, are determined variationally. The $N$ basis functions $\phi_{\mu}$, are products of radial and angular functions of hydrogenlike orbitals:

$$
\begin{gather*}
\varphi_{1 s}(1)=r_{1}^{i_{1}} e^{-\alpha r_{1}}, \quad \varphi_{1 s}(2)=r_{2}^{i_{2}} e^{-\alpha r_{2}}  \tag{3}\\
\varphi_{2 . s}(3)=\left(1-b r_{3}^{i_{3}}\right) e^{-\beta r_{3}}, \quad \varphi_{2 s}(4)=\left(1-b r_{4}^{i_{4}}\right) e^{-\beta r_{+}}  \tag{4}\\
\varphi_{2 p}(5)=r_{5}^{i_{5}} e^{-\gamma r_{5}} \cos \theta_{5} \tag{5}
\end{gather*}
$$

with $b$ as a constant to be optimized and $\mathrm{i}_{1}, i_{2}, \mathrm{i}_{3}, \mathrm{i}_{4}$, and $i_{5}$ are integers $i_{i} \geq 0$. The orbital exponents should be optimized. For the case of boron atom, with $n=5$, the unpaired electron is in a $p$-orbital, and in the non-relativistic theory, the three $p$-orbitals are degenerate in energy.

The Hylleraas ground state wave function is a function of distance coordinates of every electron and the interelectronic coordinates. In the case $n \geq 5$ the Hylleraas wave function is also function of the polar angles $\theta_{i}$ and $\phi_{i}$ of the electrons [21]. As a first step of a fullHylleraas calculation, the wave function will not yet include $r_{i j}$ terms.

The single-term wave function of five-electrons is:

$$
\begin{equation*}
\phi_{1}=e^{-\alpha r_{1}} e^{-\alpha r_{2}}\left(1-b r_{3}\right) e^{-\beta r_{3}}\left(1-b r_{4}\right) e^{-\beta r_{4}} r_{5} e^{-\gamma r_{5}} \cos \theta_{5} \tag{6}
\end{equation*}
$$

A general basis function $\phi_{\mu}$ with the powers of the $2 s$ orbital inside of the bracket, (and not outside of the bracket as would be with truly hydrogen atomic orbitals) is:

$$
\begin{equation*}
\phi_{\mu}=r_{1}^{i_{1 \mu}} e^{-\alpha r_{1}} r_{2}^{i_{2} \mu} e^{-\alpha r_{2}}\left(1-b r_{3}^{i_{3} \mu}\right) e^{-\beta r_{3}}\left(1-b r_{4}^{i_{4 \mu}}\right) e^{-\beta r_{4}} r_{5}^{i_{5 \mu}} e^{-\gamma r_{5}} \cos \theta_{5} \tag{7}
\end{equation*}
$$

where $i_{i \mu}$ are integers with $i_{i \mu} \geq 0$. The restriction for a given basis function (using for simplicity $i_{j}$ instead $i_{i \mu}$ ), $i_{1}=i_{2}$ and $i_{3}=i_{4}$, and $i_{5}$ independent of the others, ensures that the wave function is spin-eigenfunction. The wave functions (6), (7) can be evaluated into a linear combination of four basis functions constructed with Slater orbitals ${ }^{1}$ with factors in $b$. The orbital exponents $a, \beta, \gamma$ and the constant $b$ are parameters which are optimized using a simple parabolic procedure.

The nonrelativistic Hamiltonian for $n$ electrons in the field of fixed nucleus of charge $Z$ is (in a.u.):

$$
\begin{equation*}
\hat{H}=-\sum_{i=1}^{n} \frac{1}{2} \nabla_{i}^{2}-\sum_{i=1}^{n} \frac{z}{r_{i}}+\sum_{i<j}^{n} \frac{1}{r_{i j}} \tag{9}
\end{equation*}
$$

As the general Hylleraas wave function has explicit angular dependence (and also depends on $r_{i j}$ ) it is convenient to transform the kinetic energy part into mutually independent distance coordinates $r_{i}, r_{i j}$ and polar angles $\theta_{i}$ and $\varphi_{i}$. The derivation is given in detail in [21]:

$$
\begin{align*}
\hat{H} & =-\frac{1}{2} \sum_{i=1}^{n} \frac{\partial^{2}}{\partial r_{i}^{2}}-\sum_{i=1}^{n} \frac{1}{r_{i}} \frac{\partial}{\partial r_{i}}-\sum_{i=1}^{n} \frac{Z}{r_{i}}+\sum_{i<j}^{n} \frac{1}{r_{i j}}-\sum_{i<j}^{n} \frac{\partial^{2}}{\partial r_{i j}^{2}}-\sum_{i<j}^{n} \frac{2}{r_{i j}} \frac{\partial}{\partial r_{i j}}+ \\
& -\frac{1}{2} \sum_{i \neq j}^{n} \frac{r_{i}^{2}+r_{i j}^{2}-r_{j}^{2}}{r_{i} r_{i j}} \frac{\partial^{2}}{\partial r_{i} \partial r_{i j}}-\frac{1}{2} \sum_{i \neq j}^{n} \sum_{k>j}^{n} \frac{r_{i j}^{2}+r_{i k}^{2}-r_{j k}^{2}}{r_{i j} r_{i k}} \frac{\partial^{2}}{\partial r_{i j} \partial r_{i k}}+ \\
& -\frac{1}{2} \sum_{i=1}^{n} \frac{1}{r_{i}^{2}} \frac{\partial^{2}}{\partial \theta_{i}^{2}}-\frac{1}{2} \sum_{i=1}^{n} \frac{1}{r_{i}^{2} \sin ^{2} \theta_{i}} \frac{\partial^{2}}{\partial \varphi_{i}^{2}}-\frac{1}{2} \sum_{i=1}^{n} \frac{\cot \theta_{i}}{r_{i}^{2}} \frac{\partial}{\partial \theta_{i}}+  \tag{10}\\
& -\sum_{i \neq j}^{n}\left(\frac{r_{j}}{r_{i} r_{i j}} \frac{\cos \theta_{j}}{\sin \theta_{i}}+\frac{1}{2} \cot \theta_{i} \frac{r_{i j}^{2}-r_{i}^{2}-r_{j}^{2}}{r_{i}^{2} r_{i j}}\right) \frac{\partial^{2}}{\partial \theta_{i} \partial r_{i j}}+ \\
& -\sum_{i \neq j}^{n} \frac{r_{j}}{r_{i} r_{i j}} \frac{\sin \theta_{j}}{\sin \theta_{i}} \sin \left(\varphi_{i}-\varphi_{j}\right) \frac{\partial^{2}}{\partial \varphi_{i} \partial r_{i j}} .
\end{align*}
$$

[^1]\[

$$
\begin{equation*}
\phi_{\mu}=r_{1}^{i_{1} \mu} e^{-\alpha r_{1}} r_{2}^{i_{2 \mu} \mu} e^{-\alpha r_{3}} r_{3}^{i_{3} \mu} e^{-\beta r_{3}} r_{4}^{i_{4 \mu}} e^{-\beta r_{4}} r_{5}^{i_{3 \mu}} e^{-\gamma r_{3}} \cos \theta_{5} \tag{8}
\end{equation*}
$$

\]

The partial derivatives operate on the corresponding coordinates only where they appear explicitly in the wave function. For a basis function of the type of Eqs. (6), (7) which does not depend on the interelectronic coordinates, the Hamiltonian of Eq. (10) is effectively reduced to

$$
\begin{align*}
\hat{H} & =-\sum_{i=1}^{5}\left(\frac{1}{2} \frac{\partial^{2}}{\partial r_{i}^{2}}+\frac{1}{r_{i}} \frac{\partial}{\partial r_{i}}+\frac{5}{r_{i}}\right)+\sum_{i<j}^{5} \frac{1}{r_{i j}}+  \tag{11}\\
& -\frac{1}{2}\left(\frac{1}{r_{5}^{2}} \frac{\partial^{2}}{\partial \theta_{5}^{2}}+\frac{1}{r_{5}^{2}} \cot \theta_{5} \frac{\partial}{\partial \theta_{5}}\right)
\end{align*}
$$

The volume element is:

$$
\begin{equation*}
d \tau=\prod_{i=1}^{n} r_{i}^{2} \sin \theta_{i} d r_{i} d \theta_{i} d \varphi_{i} \tag{12}
\end{equation*}
$$

From the variational principle one obtains the matrix eigenvalue problem

$$
\begin{equation*}
(\mathbf{H}-E \Delta) \mathbf{C}=0 \tag{13}
\end{equation*}
$$

where the matrix elements are

$$
\begin{equation*}
H_{\mu v}=\int \psi_{\mu} \hat{H} \psi_{\nu} d \tau, \quad \Delta_{\mu v}=\int \psi_{\mu} \psi_{v} d \tau \tag{14}
\end{equation*}
$$

As the antisymmetrization operator commutes with the Hamiltonian and is idempotent, after "spin integration" we have:

$$
\begin{equation*}
H_{\mu v}=\left\langle\sum_{i=1}^{12} \phi_{\mu i}\right| \hat{H}\left|\phi_{v}\right\rangle \tag{15}
\end{equation*}
$$

On the left-hand side, the different terms are generated by the following permutations of the initial one $\phi_{\mu}$ :

$$
\begin{align*}
& \left\langle\left(\hat{e}-\hat{P}_{13}-\hat{P}_{24}+\hat{P}_{13} \hat{P}_{24}-\hat{P}_{15}-\hat{P}_{35}+\hat{P}_{15} \hat{P}_{24}+\hat{P}_{35} \hat{P}_{13}+\right.\right. \\
& \left.\left.+\hat{P}_{35} \hat{P}_{24}+\hat{P}_{35} \hat{P}_{15}-\hat{P}_{24} \hat{P}_{35} \hat{P}_{13}-\hat{P}_{24} \hat{P}_{35} \hat{P}_{15}\right) \phi_{\mu}|\hat{H}| \phi_{v}\right\rangle \tag{16}
\end{align*}
$$

where $\hat{e}$ is the unit operator and $\hat{P}_{i j}$ is the permutation operator which interchanges the spatial coordinates of electrons $i$ and $j$. If we evaluate now $H_{\mu \nu}$ we have four terms on the right hand side and $12 \times 4$ terms on the left hand side, which are products of Slater orbitals.

As in the case with Slater orbitals [21], the evaluation of the matrix element of Eq. (16) leads to the types of integrals: one-electron integrals, two-electron integrals, two-electron oneangle integrals:

$$
\begin{equation*}
\left\langle\frac{\cos \theta_{i}}{r_{i j}}\right\rangle=0, \quad\left\langle\frac{\cos ^{2} \theta_{i}}{r_{i j}}\right\rangle=0, \quad \forall_{i<j} \tag{17}
\end{equation*}
$$

and two-electron two-angle integrals:

$$
\begin{equation*}
\left\langle\frac{\cos \theta_{i} \cos \theta_{j}}{r_{i j}}\right\rangle=\frac{(4 \pi)^{2}}{9}\left\langle\frac{s_{i j}}{g_{i j}^{2}}\right\rangle, \quad \forall_{1<i} \tag{18}
\end{equation*}
$$

where $s_{i j}$ is the smallest of $r_{i}$ and $r_{j}$, and $g_{i j}$ the largest of $r_{i}$ and $r_{j}$.
The resulting integrals in the Hylleraas method have been developed into programmable expressions using the MAPLE [22] program package. The one-electron and two-electron Slater integrals are solved with the help of auxiliary integrals [9]. They are once calculated and stored, and read when it is necessary.

The optimization procedure is simple. The three orbital exponents and the constant $b$ are optimized one after the other using a numerical one-dimensional procedure. A three-point parabolic fitting is mixed with two other procedures which are the selection of the point with lower energy and the best virial coefficient. Each variable is changed with a different step-size which decreases with the cycles of optimization. For each energy calculation the virial coefficient is calculated to check the accuracy of the calculation.

## 3. CALCULATIONS

The first basis function $\phi_{1}$, Eq. (6), can be considered as an approximation to the HartreeFock wave function constructed with medium size basis set of orbitals.

Then the energy is given by

$$
\begin{equation*}
E_{1}=\frac{\left\langle\psi_{1}\right| \hat{H}\left|\psi_{1}\right\rangle}{\left\langle\psi_{1} \mid \psi_{1}\right\rangle} \tag{19}
\end{equation*}
$$

where $\psi_{1}=\hat{A} \phi_{1} \chi$. We need a condition to calculate the energy of the state. The condition is that the virial theorem should be fullfiled. Therefore a factor $\eta$, the ratio between potential $\langle V\rangle$ and kinetic energies $\langle T\rangle$ should be precisely $2 \pm 10{ }^{10}$. This factor $\eta$ is evaluated in every calculation to ensure that the state is the appropriate one:

$$
\begin{equation*}
\eta=-\frac{\langle V\rangle}{\langle T\rangle} \tag{20}
\end{equation*}
$$

The optimization of the exponents $\alpha, \beta, \gamma$ and the constant $b$ leads to the energy $E_{1}=24.501187$ a.u., virial factor $\eta=1.9999999999, \alpha=4.688051, \beta=1.357372 \gamma=1.204539$ and $b=6.5242$. This energy is lower than the energy the calculated with Slater orbitals [21] -24.498369 a.u. The optimized exponents $\alpha, \gamma$, of the wave function with hydrogen-like
orbitals are very close to those obtained using the wave function with Slater orbitals, $\alpha=$ 4.6794196, $\gamma=1.2106724$. The reason is that the $1 s$ and $2 p$ orbitals have the same form in both wave functions and they are only weakly dependent on the form of the $2 s$ orbital. As it is expected, the exponent of the $2 s$ orbital is different compared with $\beta=1.2880853$ for Slater orbitals. In the optimization of $b$ one finds a very flat parabola.

The energy lies between the Hartree-Fock energy - 24.498369 a.u. calculated using a single-zeta basis set and -24.527920 a.u. obtained with a double-zeta basis set [23]. Other results are given in Table IV. The energy obtained using hydrogenlike orbitals is lower than using Slater orbitals.

In this method the energy is improved by adding new basis functions in a systematic way, from $i_{j}=0$ until $i_{j}=7$, in some cases $i_{j}=9$, ordered by similarity to facilitate the optimization of the exponents. The orbital exponents and the constant $b$ have been optimized for every group of similar basis functions which is added. We have used the same basis function expansion than in the previous calculations with Slater orbitals [21]. In Table 1, the powers of the orbitals of the basis functions are given. For example, the basis functions 1 to 38 introduce flexibility to the outer electrons. The basis functions 39 to 57 give flexibility to the inner electrons. The rest of the basis functions introduces some correlation effects due to the use of more extended orbitals.

The energy results are shown in Table 2, together with the energies of those basis function expansions using Slater orbitals. For the first eight basis functions we have $i_{3}, i_{4}=0$, as the basis set is the same, the form of the hydrogenlike orbital $2 s$ coincides with the Slater one for these basis functions except for a factor, therefore the energy results for both wave functions are equal. Taking more basis functions, the wave function of hydrogenic orbitals lead to an energy about 1 mhartree lower than the energy calculated with Slater orbitals. The difference increases with increasing number up to 150 basis functions. The final energy for $N=150$ is $\mathrm{E}_{150}=-24.550233$ a.u., 8.98 mhartree lower than the energy with Slater basis functions. This energy includes about 33 per cent of the total correlation energy for boron atom. 6 per cent more correlation energy have been obtained using hydrogenlike orbitals instead of Slater ones. This is due that they are physically more appropiate and to the presence of one more variational parameter, the constant $b$.

The optimized exponents and constants are given in Table 3. The final exponents are $\alpha=4.520884, \beta=1.338861$, and $\gamma=0.750846$, and the constant is fitted to $b=2.398$. These exponents are slightly smaller than the ones calculated using Slater orbitals. In the first basis functions $\alpha$ and $\gamma$ coincide with the ones of Slater, but not $\beta$. Adding more basis functions the exponents became more different. In contrary, the $\beta$ exponent became more close to the Slater one.

In Figure 1 the radial distribution functions of the Slater and hydrogenic $2 s$ orbital using the optimized parameters of the 150 -term wave function are shown. The hydrogenic orbital

Table 1. Basis function powers: $\phi_{\mu}=r_{1}^{i_{1}} r_{2}^{l_{2}}\left(1-b r_{3}^{i_{3}}\right)\left(1-b r_{4}^{I_{4}}\right) r_{5}^{i_{5}} \cos \theta_{5} e^{\left(-\alpha r_{1}-\alpha r_{2}-\beta r_{3}-\beta r_{4}-\gamma r_{3}\right)}$

| No. | $i_{1}, i_{2}$ | $i_{3}, i_{4}$ | $i_{5}$ | No | $i_{1}, i_{2}$ | $i_{3}, i_{4}$ | $i_{5}$ | No. | $i_{1}, i_{2}$ | $i_{3}, i_{4}$ | $i_{5}$ | No. | $i_{1}, i_{2}$ | $i_{3}, i_{4}$ | $i_{5}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1 | 0 | 0 | 0 | 39 | I | 0 | 0 | 77 | 3 | 4 | 1 | 115 | 4 | 4 | 2 |
| 2 | 0 | 0 | 1 | 40 | 2 | 0 | 0 | 78 | 4 | 2 | 1 | 116 | 4 | 4 | 3 |
| 3 | 0 | 0 | 2 | 41 | 3 | 0 | 0 | 79 | 4 | 3 | 1 | 117 | 4 | 4 | 4 |
| 4 | 0 | 0 | 3 | 42 | 4 | 0 | 0 | 80 | 1 | 2 | 2 | 118 | 0 | 0 | 8 |
| 5 | 0 | 0 | 4 | 43 | 1 | 1 | 0 | 81 | 1 | 2 | 3 | 119 | 0 | 0 | 9 |
| 6 | 0 | 0 | 5 | 44 | 2 | 1 | 0 | 82 | 1 | 2 | 4 | 120 | 1 | 5 | 0 |
| 7 | 0 | 0 | 6 | 45 | 3 | 1 | 0 | 83 | 1 | 2 | 5 | 121 | 1 | 6 | 0 |
| 8 | 0 | 0 | 7 | 46 | 4 | 1 | 0 | 84 | 1 | 3 | 2 | 122 | 2 | 5 | 0 |
| 9 | 0 | 1 | 0 | 47 | 1 | 2 | 0 | 85 | 1 | 3 | 3 | 123 | 2 | 6 | 0 |
| 10 | 0 | 2 | 0 | 48 | 1 | 3 | 0 | 86 | 1 | 3 | 4 | 124 | 3 | 5 | 0 |
| 11 | 0 | 3 | 0 | 49 | 1 | 4 | 0 | 87 | 1 | 4 | 4 | 125 | 3 | 6 | 0 |
| 12 | 0 | 4 | 0 | 50 | 2 | 2 | 0 | 88 | 2 | 2 | 2 | 126 | 4 | 4 | 0 |
| 13 | 0 | 5 | 0 | 51 | 2 | 3 | 0 | 89 | 2 | 2 | 3 | 127 | 4 | 5 | 0 |
| 14 | 0 | 1 | 1 | 52 | 2 | 4 | 0 | 90 | 2 | 2 | 4 | 128 | 1 | 5 | 1 |
| 15 | 0 | 1 | 2 | 53 | 3 | 2 | 0 | 91 | 2 | 2 | 5 | 129 | 1 | 6 | 1 |
| 16 | 0 | 1 | 3 | 54 | 3 | 3 | 0 | 92 | 2 | 3 | 2 | 130 | 2 | 5 | 1 |
| 17 | 0 | 1 | 4 | 55 | 3 | 4 | 0 | 93 | 2 | 3 | 3 | 131 | 2 | 6 | 1 |
| 18 | 0 | 1 | 5 | 56 | 4 | 2 | 0 | 94 | 2 | 3 | 4 | 132 | 3 | 5 | 1 |
| 19 | 0 | 1 | 6 | 57 | 4 | 3 | 0 | 95 | 2 | 3 | 5 | 133 | 3 | 6 | 1 |
| 20 | 0 | 1 | 7 | 58 | 1 | 1 | 1 | 96 | 3 | 3 | 2 | 134 | 4 | 5 | 1 |
| 21 | 0 | 2 | 1 | 59 | 1 | 1 | 2 | 97 | 3 | 3 | 3 | 135 | 4 | 5 | 2 |
| 22 | 0 | 2 | 2 | 60 | 1 | 1 | 3 | 98 | 3 | 3 | 4 | 136 | 4 | 5 | 3 |
| 23 | 0 | 2 | 3 | 61 | 1 | 1 | 4 | 99 | 3 | 3 | 5 | 137 | 5 | 4 | 1 |
| 24 | 0 | 2 | 4 | 62 | 1 | 1 | 5 | 100 | 3 | 2 | 2 | 138 | 5 | 4 | 2 |
| 25 | 0 | 2 | 5 | 63 | 1 | 1 | 6 | 101 | 3 | 2 | 3 | 139 | 5 | 4 | 3 |
| 26 | 0 | 2 | 6 | 64 | 1 | 1 | 7 | 102 | 3 | 2 | 4 | 140 | 5 | 4 | 4 |
| 27 | 0 | 3 | 1 | 65 | 1 | 1 | 8 | 103 | 3 | 2 | 5 | 141 | 5 | 5 | 0 |
| 28 | 0 | 3 | 2 | 66 | 2 | 1 | 1 | 104 | 2 | 4 | 2 | 142 | 5 | 5 | 1 |
| 29 | 0 | 3 | 3 | 67 | 3 | 1 | 1 | 105 | 2 | 4 | 3 | 143 | 5 | 5 | 2 |
| 30 | 0 | 3 | 4 | 68 | 4 | 1 | 1 | 106 | 2 | 4 | 4 | 144 | 5 | 5 | 3 |
| 31 | 0 | 3 | 5 | 69 | 1 | 2 | 1 | 107 | 3 | 4 | 2 | 145 | 1 | 5 | 2 |
| 32 | 0 | 4 | 1 | 70 | 1 | 3 | 1 | 108 | 3 | 4 | 3 | 146 | 1 | 6 | 2 |
| 33 | 0 | 4 | 2 | 71 | 1 | 4 | 1 | 109 | 3 | 4 | 4 | 147 | 2 | 5 | 2 |
| 34 | 0 | 4 | 3 | 72 | 2 | 2 | 1 | 110 | 4 | 3 | 2 | 148 | 2 | 6 | 2 |
| 35 | 0 | 4 | 4 | 73 | 2 | 3 | 1 | 111 | 4 | 3 | 3 | 149 | 3 | 5 | 2 |
| 36 | 0 | 5 | 1 | 74 | 2 | 4 | 1 | 112 | 4 | 3 | 4 | 150 | 3 | 6 | 2 |
| 37 | 0 | 5 | 2 | 75 | 3 | 2 | 1 | 113 | 4 | 3 | 5 |  |  |  |  |
| 38 | 0 | 5 | 3 | 76 | 3 | 3 | 1 | 114 | 4 | 4 | 1 |  |  |  |  |

Table 2. Energies (in a.u.) for a given number of basis functions, using Slater orbitals [21], and hydrogenlike orbitals in the wave function, respectively. In addition the difference $E$ (hydrogenlike) $E$ (Slater) is given

| No. | $E$ (Slater) | $E$ <br> (hydrogenlike) | Difference | No. | $E$ (Slater) | $E$ <br> (hydrogenlike) | Difference |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1 | -24.348104 | -24.351385 | -0.003281 | 77 | -24.540131 | -24.542418 | -0.002287 |
| 8 | -24.432419 | -24.432427 | -0.000008 | 79 | -24.540146 | -24.542464 | -0.002318 |
| 13 | -24.491764 | -24.494885 | -0.003121 | 87 | -24.540156 | -24.542582 | -0.002426 |
| 20 | -24.516098 | -24.516143 | -0.000045 | 91 | -24.540246 | -24.542602 | -0.002356 |
| 26 | -24.518749 | -24.518898 | -0.000149 | 95 | -24.540293 | -24.542609 | -0.002316 |
| 31 | -24.519192 | -24.519968 | -0.000776 | 99 | -24.540292 | -24.542609 | -0.002317 |
| 35 | -24.519215 | -24.520470 | -0.001255 | 103 | -24.540305 | -24.542613 | -0.002308 |
| 38 | -24.519225 | -24.520575 | -0.001350 | 109 | -24.540309 | -24.542621 | -0.002312 |
| 42 | -24.525574 | -24.528064 | -0.002490 | 117 | -24.540337 | -24.542628 | -0.002291 |
| 46 | -24.535417 | -24.536589 | -0.001172 | 127 | -24.540697 | -24.545418 | -0.004721 |
| 49 | -24.536008 | -24.537109 | -0.001101 | 129 | -24.540700 | -24.545960 | -0.005260 |
| 52 | -24.536502 | -24.538329 | -0.001827 | 131 | -24.540719 | -24.545987 | -0.005268 |
| 55 | -24.536553 | -24.538352 | -0.001799 | 136 | -24.540730 | -24.545990 | -0.005260 |
| 57 | -24.536607 | -24.538491 | -0.001884 | 137 | -24.540762 | -24.546496 | -0.005734 |
| 65 | -24.539255 | -24.540783 | -0.001528 | 140 | -24.540753 | -24.547476 | -0.006723 |
| 68 | -24.539624 | -24.540786 | -0.001162 | 144 | -24.541177 | -24.549774 | -0.008597 |
| 71 | -24.539974 | -24.542265 | -0.002291 | 150 | -24.541246 | -24.550233 | -0.008987 |
| 74 | -24.540018 | -24.542364 | -0.002346 |  |  |  |  |



Fig. 1. Radial distribution function of the $2 s$ hydrogenlike orbital (with a node) and the Slater type one calculated with the optimized exponents of a 150 -term wave function
has a node and the maximal probability at larger distance. For $i_{3}, i_{4}=1$ the hydrogenlike orbital coincides with the hydrogenic orbital.

In Figures 2 and 3 the radial distribution functions of the 2 s orbitals with different powers $i_{3}, i_{4}$ show different behavior at small $r$. The hydrogenlike orbitals present nodes, which move to large with increasing powers. The Slater orbitals have a maximum at nearer distances to
the nucleus than the hydrogenlike ones. In both Figures the probability to find the electron is higher for power 1 and decreases with increasing power.

Table 3. Results using hydrogenlike orbitals: energies in (a. u.) for a given number of basis functions, optimized orbital exponents and orbital coefficient $b$, and virial coefficient

| No. | (hydrogenlike) | $\alpha$ | $\beta$ | $\gamma$ | $b$ | virial ratio |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: |
| 1 | -24.351385 | 4.62183388 | 0.74896367 | 0.47489419 | 7.136869 | 2.0000000344 |
| 8 | -24.432427 | 4.62670884 | 0.72026697 | 1.65453049 | 5.852452 | 2.0000080293 |
| 13 | -24.494885 | 4.67718391 | 1.05324185 | 0.64033174 | 2.753014 | 2.0000000000 |
| 20 | -24.516143 | 4.68590000 | 1.24058271 | 0.63399000 | 4.599031 | 2.0000000000 |
| 26 | -24.518898 | 4.68620737 | 1.27517474 | 0.88807764 | 6.834067 | 2.0000000000 |
| 31 | -24.519968 | 4.68641492 | 1.24694281 | 1.60862117 | 5.470780 | 2.0000000000 |
| 35 | -24.520470 | 4.68641492 | 1.22935089 | 1.61343967 | 4.077469 | 2.0000000000 |
| 38 | -24.520575 | 4.68676343 | 1.21088914 | 1.61937951 | 3.785341 | 1.9999999971 |
| 42 | -24.528064 | 4.55953982 | 1.23382125 | 0.72695372 | 2.919047 | 2.0000000000 |
| 46 | -24.536589 | 4.45773982 | 1.25648021 | 0.69123238 | 3.517716 | 2.0000000000 |
| 49 | -24.537109 | 4.44461335 | 1.23386413 | 0.68764275 | 3.571725 | 2.0000000000 |
| 52 | -24.538329 | 4.44841053 | 1.21365925 | 0.69109923 | 3.030002 | 2.0000000000 |
| 55 | -24.538352 | 4.44975485 | 1.20919613 | 0.69322458 | 2.948800 | 1.9999999034 |
| 57 | -24.538491 | 4.44701898 | 1.20534653 | 0.69306467 | 2.949590 | 1.9999998022 |
| 65 | -24.540784 | 4.42490154 | 1.20982109 | 0.74509752 | 3.108249 | 2.0000000015 |
| 68 | -24.540786 | 4.42490154 | 1.21082109 | 0.74841752 | 3.125801 | 2.00000000015 |
| 71 | -24.542265 | 4.41928966 | 1.21987149 | 1.00909552 | 2.872800 | 1.9999999081 |
| 74 | -24.542364 | 4.41762286 | 1.21987149 | 1.00909552 | 2.872800 | 1.9999999131 |
| 77 | -24.542418 | 4.41653618 | 1.21987149 | 1.00909552 | 2.867022 | 1.9999999894 |
| 79 | -24.542464 | 4.41660018 | 1.21987149 | 1.00909552 | 2.855222 | 2.0000000509 |
| 87 | -24.542582 | 4.42267218 | 1.21987149 | 1.00909552 | 2.783861 | 1.9999999919 |
| 91 | -24.542602 | 4.42237993 | 1.21987149 | 1.00909552 | 2.780100 | 1.9999999989 |
| 95 | -24.542609 | 4.42239693 | 1.21987149 | 1.00909552 | 2.778428 | 1.9999999999 |
| 99 | -24.542609 | 4.42226869 | 1.21986069 | 1.00909732 | 2.779800 | 1.9999999917 |
| 103 | -24.542613 | 4.42214949 | 1.21986133 | 1.00908742 | 2.779000 | 1.9999999980 |
| 109 | -24.542621 | 4.42229789 | 1.21984899 | 1.00898732 | 2.771000 | 1.9999999973 |
| 117 | -24.542628 | 4.42206969 | 1.21985069 | 1.00909738 | 2.778000 | 1.9999996485 |
| 127 | -24.545419 | 4.39900000 | 1.27000160 | 0.83784000 | 2.599200 | 1.9999999171 |
| 129 | -24.545960 | 4.40000000 | 1.26016160 | 0.89900000 | 2.498000 | 1.9999999659 |
| 131 | -24.545987 | 4.40101164 | 1.27035180 | 0.91071746 | 2.509400 | 2.0000001621 |
| 136 | -24.545990 | 4.40116722 | 1.28307195 | 0.92082000 | 2.504400 | 1.9999998322 |
| 137 | -24.546496 | 4.40426969 | 1.23087969 | 0.89209732 | 2.578000 | 2.0000000398 |
| 140 | -24.547476 | 4.41000000 | 1.24979687 | 0.88833184 | 2.388000 | 2.00000005333 |
|  | -24.549774 | 4.51281980 | 1.33886068 | 0.71579732 | 2.398000 | 2.0000000454 |
|  | -24.550233 | 4.52088419 | 1.33886068 | 0.75084613 | 2.398000 | 2.0000000481 |
|  |  |  |  |  |  |  |



Fig. 2. Radial distribution functions of the $2 s$ hydrogenlike orbitals with different powers of $r$, $i_{3}=i_{4}$ from Eq. (7). Calculated with the optimized exponents of the 150 -term wave function

Fig. 3. Radial distribution functions of the $2 s$ Slater orbitals with different powers in $r, i_{3}=i_{4}$ from Eq. (8). Calculated with the optimized exponents of the 150 -term wave function

Table 4. Comparison of the energies for boron atom calculated with different methods and the nonrelativistic energy (all values in a.u.)

| Reference | Year | Method | Basis Set | Energy |
| :--- | :---: | :--- | :--- | :--- |
| Clementi [23] | 1989 | HF | c-GTO small geom. | -24.495670 |
| Clementi, Roetti [24] | 1974 | HF | STO single zeta | -24.498369 |
| Ruiz [21] | 2004 | ref Hy | (1-term) | -24.498369 |
| Present Work | 2004 | ref Hy hydrog. | (1-term) | $-\mathbf{- 2 4 . 5 0 1 1 8 7}$ |
| Clementi, Raimondi [25] | 1963 | HF | STO minimal basis | -24.498370 |
| Clementi, Roetti [24] | 1974 | HF | STO double zeta | -24.527920 |
| Clementi [23] | 1989 | HF | GTO large geom. | -24.528486 |
| Huzinaga [26] | 1977 | HF | STO | -24.528709 |
| Froese-Fisher et La [29] | 1993 | numerical HF |  | -24.529036 |
| Clementi, Roetti [24] | 1974 | HF | STO extended | -24.529057 |
| Clementi, Chakravorty [23] | 1989 | HF | GTO extra large | -24.52906 |
|  | 2004 | HF-Gaussian-98 | cc-pVSZ | -24.5291094 |
| Mayer [28] | 2004 | full-CI | 4-31G | -24.530874 |
| Ruiz [21] | 2004 | Hylleraas | I50 (without $r_{i j}$ ) | -24.541246 |
| Present Work | 2004 | Hylleraas | 150 (without $r_{i j}$ ) | -24.550233 |
| Froese-Fisher et la [30] | 1994 | multiref. CI |  | -24.560354 |
| Mayer [28] | 2004 | full-CI | DZV | -24.568999 |
| Mayer [28] | 2004 | full--CI | DZP | -24.585533 |
| Estimated nonrelativistic [31] | 1993 |  | -24.65391 |  |

Small geom. stands for small geometrical basis set, large geom. for large geometrical basis set, DZV for valencc double zeta basis and DZP for double zeta basis with polarization functions. In the literature, usually the best HF energy value is quoted as only HF, this is -24.52905 a.u.

Finally, in Table 4 we compare our energy results with the ones of other methods. The 150 -term energy obtained using hydrogenlike orbitals is lower by 19.3 mhartrees than the energy of a full-CI wave function with $4-31 \mathrm{G}$ basis set [28]. It is also lower than numerical HF calculations. The calculated energy is higher by 10.1 mhartrees than the multireference CI energy [29], and 18.7 mhartrees higher than the full-CI energy using a double-zeta basis set. The energy using polarization basis functions is still lower. A similar energy to the full-CI using medium size basis set (DZV) could probably be achieved if the expansion serie would be extended to a larger number of basis functions. Nevertheless, our purpose is to introduce more correlation energy introducing interelectronic coordinates into the wave function, this work is in progress.

## 4. CONCLUSIONS

Substituting the $2 s$ Slater orbital by the hydrogenlike $2 s$ one into the reference Hylleraas wave function the energy result improves by 6 per cent for a large number of basis functions. This is due, on the one hand, to the higher quality of the $2 s$ orbital, which has a node, as the exact solution of Schrödinger equation of the hydrogen atom. On the other hand, it is due to the presence of an aditional variational parameter. The orbitals with high powers in $r$ of the expansion series are not exactly Slater neither hydrogenic, they are of $s$-type. It can be understood as a basis function set inside of a shell. The obtained energy approaches the full-CI calculations using valence double-zeta basis set. We expect to introduce more correlation effects by introducing in this reference wave function interelectronic variables. This work is in progress.

Finally, the calculated energy is comparable with the energy of full-CI calculations.

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[^0]:    * Dedicated to the memory of Professor Jacek Rychlewski

[^1]:    ${ }^{1}$ This fact makes the calculation of matrix elements more cumbersome as if one uses the basis function with Slater orbitals [21]

